Principles of data reduction: infrared imaging

- NOTCam survival guide -

Erkki Kankare 15th October 2013 FINCA observing school











Near-infrared (near-IR)



- ~ 7000 Å 5 μm (not visible for human eye)
- Atmosphere mostly transparent in near-IR (ground-based telescopes useful)

Nordic Optical Telescope near-infrared Camera and spectrograph (NOTCam)

- www.not.iac.es/instruments/notcam/
- Rockwell "HAWAII" HgCdTe array
- 1024x1024x18.5µm pixels
- Wide-field imaging
 - 4' x 4' (0.234"/pixel)
- High-resolution imaging
 - 80" x 80" (0.078"/pixel)
- Long-slit spectroscopy
- 0.8 2.5 μm wavelength range
- 4 quadrants in the array



- Fixed setup
 - Broadband imaging JHKs
 - Selection of other filters
- All filters have their own focus offset (foc-del) compared to the WF K-band

Im	Imaging mode focus offsets found in good seeing (fwhm:0.3"-0.5").							
	WF Camera	HR Camera						

	WF Camera	HR Camera	
	5650	20	Internal camera focus (fixed)
foc-pos	23450 *)	23450 *)	Telescope value for K,K',Ks
foc-del	0	110	K, K', Ks
foc-del	-45	70	Н
foc-del	-100	30	J
foc-del	-160	-50	Y
foc-del	-200	-75	Z

www.not.iac.es/instruments/notcam/filters/index.html



NOTCam	status	Name	Bandpass (µm))	Transmission	Comment
Filter #	Status	(click for details)	λου	λ central	λ _{off}	Curves	Comment
201	Mounted	J	1.165		1.328	<u>gif,ascii</u>	NOTCam standard J
202	Mounted	<u>J</u> (spare)	1.167	1.250	1.334	<u>gif,ascii</u>	Mounted in the stop wheel
203	Mounted	<u>H</u>	1.484		1.780	<u>gif,ascii</u>	NOTCam standard H
204	Mounted	<u>H</u> (spare)	1.487	1.626	1.766	<u>gif,ascii</u>	Mounted in the stop wheel
205	Mounted	<u>K'</u>	1.950		2.290	gif,ascii	
206	Mounted	<u>K'</u> (spare)	1.950	2.115	2.280	<u>gif,ascii</u>	Mounted in the stop wheel
207	Mounted	Ks	1.999	2.140	2.282	<u>gif,ascii</u>	NOTCam standard K for imaging
208	Mounted	K	2.038	2.200	2.363	<u>gif,ascii</u>	NOTCam standard K for spectroscopy
209	Mounted	Br Y	2.147	2.163	2.179	<u>ps,ascii</u>	
210	Mounted	K continuum	2.251	2.267	2.284	<u>ps,ascii</u>	
211	Mounted	H continuum	1.562	1.574	1.586	<u>ps,ascii</u>	
212	Mounted	[Fe II]	1.632	1.645	1.657	<u>ps,ascii</u> ,	
213	Mounted	He I _A	1.070	1.079	1.089	png.	
214	Mounted	Раў	1.086	1.094	1.102	•	
215	Mounted	J-continuum	1.201	1.211	1.221	<u>ps,ascii</u>	
216	Mounted	Paβ	1.277	1.287	1.296	<u>ps,ascii</u>	
217	Mounted	He I _B	2.056	2.071	2.086	<u>ps,ascii</u>	
218	Mounted	H ₂ v=1-0 S(1)	2.101	2.118	2.133	<u>ps,ascii</u>	
219	Mounted	He I _C	2.167	2.184	2.200	<u>ps,ascii</u>	
220	Mounted	H ₂ v=2-1 S(1)	2.233	2.251	2.268	<u>ps,ascii</u>	
221	Mounted	CO (2-0 bandhead)	2.271	2.288	2.305	<u>ps,ascii</u>	
222	Mounted	Yn	1.004	1.028	1.053	png	Check note!
223	Mounted	CH ₄ s	1.543	1.599	1.655	<u>ps,ascii</u>	
224	Mounted	CH ₄ I	1.624	1.680	1.736	<u>ps,ascii</u>	
225	Order cancelled	Grism 1	1.000	1.300	1.600		
226	Stored	Grism 2	1.400	1.950	2.500	1	
227	Stored	[Fe II]	×	1.644			From Barr, use #212.
228	Mounted	[Fe II] Continuum	1.682	1.689	1.696	<u>ps,ascii</u>	From Barr.
229	Stored	H ₂ v=1-0 S(1)		2.122		10	From Barr, use #218.
230	Mounted	H ₂ v=1-0 S(1) Continuum	2.077	2.087	2.097	<u>ps,ascii</u>	From Barr.
231	Stored	1754/10		1.754			Belongs to B. Thomsen
232	Returned	1004/10		1.004		2	Belongs to G. Östlin
233	Returned	1012/10		1.012			Belongs to G. Östlin
234	Returned	1029/10		1.029			Belongs to G. Östlin
235	Returned	1020/10	×	1.020			Belongs to G. Östlin
236	Mounted	Ϋ́	0.944	1.020	1.097	png1, png2, png3, ascii	From NDC Sep-2010.
237	Mounted	Z	0.830	0.889	0.949	jpg1, jpg2, jpg3, <u>ascii</u>	From NDC Dec-2010.
238	Mounted	<u>BK7</u>	-	-	-	jpg, ascii	From Custom Scientific Mar-2012.
239	Mounted	KG4	-	1946	-	jpg, <u>ascii</u>	From Custom Scientific Mar-2012.

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	Filter	#209 Broamma vacuum waveleoot	ь т_74к	4 574	1 500	<u>ps,ascii</u>	
_ 100		#203, Di gamina, vacuum wavelengu	.1, 1-741			<u>ps,ascii,</u>	
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							Belongs to G. Östlin
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2110	2120 2130 21	40 2150 2160 2170	2180	2190	2200 2	210 jpg1, jpg2, jpg3, ascii	From NDC Dec-2010.
		wavelengin (nm)	1	I		jpg, ascii	From Custom Scientific Mar-2012.
239	Mounted	KG4	-	200	-	jpg, <u>ascii</u>	From Custom Scientific Mar-2012.



Bad pixels & Calibration

- www.not.iac.es/instruments/qc/
- Dead column in the centre of the array due to the readout 'feature'.
 - Also many cold pixels
- The only calibrations required for NOTCam near-IR imaging are differential (bright and faint) sky flats. Faint frames are used to subtract the thermal emission (independent of the sky intensity) from the bright frames.
- No bias subtraction
- Darks can be used to make a bad pixel mask (this is also available on the NOT web pages)
- No dark subtraction NOTCam dark current is un-stable (and sky subtraction is done anyway)
- Dark exposures can be used to reset memory effects on a saturated array
 - clean3





Distortion

- www.not.iac.es/instruments/notcam/calibration.html
- WF camera suffers from optical distortion
- Distortion correction model available
- Uncorrected (left) and distortion corrected (right) Hband images of the field used to make the model





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GAIN & RON



- www.not.iac.es/instruments/qc/
- Gain, relation between photons/electrons and ADUs/counts: e⁻/ADU (ADU = Analogue-to-Digital Unit)
- Ramp-sampling mode (frame mode), gain ~ 2.6 e⁻/ADU

GAIN & RON



- www.not.iac.es/instruments/qc/
- RON (read-out-noise)
- Ramp-sampling mode (frame mode), RON ~ 8.6 e⁻

Linearity



- www.not.iac.es/instruments/qc/
- Linearity region: range of counts with a constant gain within a small (<1%) error</p>
 - Typical problem with near-IR arrays (CCDs linear up to the saturation limit)
- Linearity <25000 counts</p>

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Sky background

- In near-IR the sky background is always high
 - Typical exposure of 6x10 sec with frame mode
 - In J-band ~3000-5000 ADU
 - In H-band ~15000-25000 ADU
 - In Ks-band ~20000-30000 ADU
- Sky subtraction crucial!
 - For compact sources dithered (jittered) target frames as sky frames
 - For extended sources separate sky frames (beam-switch mode)
 - Target and sky observations require the same total exposure time
- Observing strategy
 - Sky changes rapidly and only <1-2 min exposures should be used</p>
 - Pay attention to the linearity!!
 - Adjust exposure parameters for the sequence, if necessary
 - Multiple short exposures with a dither pattern (5-point dice, 3x3 grid)
 - Repeat sequence if necessary for the desired S/N ratio
 - Ist exposure with a new filter usually low-quality

Exposure commands: exp, mexp, frame

- Reset-read-read mode (Fowler sampling):
 - reset the array, read out a reset frame, integrate, read-out, on-line subtract reset frame.
 - Used by exp t (single t sec exposure) and mexp t N (N exposures of t sec, the average of which is the final image).
 - Image count level t sec
 - Final single image *.fits[1], reset frame *.fits[2]
 - mexp creates a separate average file to use (extension *fits[1])
- Ramp-sampling mode:
 - reset the array, read out a reset frame, integrate, multiple read-outs during the integration.
 - Used by frame t N (t*N exposure corrected with linear regression analysis to reduce the noise by a factor of N^{1/2})
 - Image count level t * N sec
 - Final image *.fits[1], read-out files *fits[2] ... *fits[N+1], reset frame *fits[N+2].
- NOTCam files, like all NOT files, are Multiple Extensions FITS (MEF) files
- Integration time + Overheads = 1.4 x integration time (more with beamswitch)

Exposure commands: frame

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 Vertical pattern is called a "reset anomaly" ("bias tilt", "dc-gradient"), typical for Hawaii arrays. Present independent of the used exposure modes.

Exposure commands: mexp

- Look for an empty EXPMODE keyword in the header for the average files
- Average files are also byte-wise half the size of the original files
- Example: 5 (dice) J-band images observed with mexp 15 4
 - Files needed for reduction: NCve060274.fits , NCve060279.fits , NCve060284.fits , NCve060289.fits , NCve060294.fits

xankare-latitude: ~/Arp299_20120506
Search Terminal Help its Arp299_J 5 mexp 15.0 4 3/4
its Arp299_J 5 mexp 15.0 4 3/4
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re-latitude:~/Arp299_20120506\$

Field orientation

- Default field orientation of NOTCam is -90
 - North is Left, East is Up
- field-rot = -90 to minimize the spilling of LN_2



Beam-switch & Guiding

- Separate sky frames for extended targets
- Avoid bright stars on the OFF (sky) field
- Figure on the right: Example of beamswitching with NOTCAM-N-negX
- Beamswitch script ON-OFF sequence: target – sky – target – sky …

Below is a listing of all available NOTCam guide areas:

- NOTCAM-De NOTCam_Default field (60" step guaranteed)
- NOTCAM-Ma NOTCam_Maximum field (not useful for dithering)
- NOTCAM-Re NOTCam_Reduced field (100" step guaranteed)
- NOTCAM-Be-NS Beamswitch offset in NS direction
- NOTCAM-Be-EW Beamswitch offset in EW direction
- NOTCAM-S-posX Beamswitch offset to the South (target moves in positive X direction)
- NOTCAM-N-negX Beamswitch offset to the North (target moves in negative X direction)
- NOTCAM-E-negY Beamswitch offset to the East (target moves in negative Y direction)
- NOTCAM-W-posY Beamswitch offset to the West (target moves in positive Y direction)



NOTCam observer's graphical user interface



- Optical elements on the light path listed
- GUI to manually select optical elements
- 2 x ds9
- Sequencers for writing commands
- Status windows for the instrument/telescope
- Remote observers have a simplified version



How to observe with NOTCam?

- www.not.iac.es/observing/seq/notcam-seq-scripts.html
- www.not.iac.es/observing/seq/notcam-seq-commands.html
- www.not.iac.es/observing/forms/signal/v2.2/index.php
- Key notcam script 1: Setup imaging camera mode, filter and filter focus offset
 - setup-ima camera filter-ID
 - Example, setup H (#203) band with WF camera:
 - setup-ima WF 203
- Key notcam script 2: Sequence of observations with a 3x3 grid
 - 9point rmode t N "obj" step skew Ncycles
 - Example, 1x9-point with 6x5sec per pointing in frame mode of SN 2012ej with 40" dither step and a 2" skew:
 - 9point frame 5 6 SN2012ej 40 2 1
- Key notcam script 3: Beam-switch (takes 3x3 grid sequence both ON and OFF)
 - beamswitch rmode t N "obj" dir beam step skew
 - Example: 1x9-point with 6x5sec per pointing in frame mode of SN 2012ej with 10" dither step, a 2" skew and separate sky:
 - beamswitch frame 5 6 SN2012ej N-negX 180 10 2

notcam package installation

- Installation instructions: www.not.iac.es/instruments/notcam/quicklook.README
- To summarise, open a new terminal and type:
 - \$ cd
 - \$ cd iraf
 - \$ mkdir extern
 - \$ cd extern
 - \$ wget www.not.iac.es/instruments/notcam/quicklook.tar.gz
 - \$ tar xzf quicklook.tar.gz
 - \$ cd ..
 - \$ emacs login.cl
 - Add the following lines after the first 'keep' line
 - reset notcam = /home/myusername/iraf/extern/notcam_v2.5/
 - task notcam.pkg = notcam\$notcam.cl
 - printf ("reset helpdb=%s,notcam\$lib/helpdb.mip\nkeep\n",
 - envget("helpdb")) | cl
 - Save and exit

Reduction example 1: files & lists

- Files needed for reduction: flats and target frames
- gethead command useful in reading file headers quickly
- Example data available in /course_data/nir/SN2012ej_20120907.tar.gz
- Create lists of files which include all the file names of 1 target in 1 band
- Create lists of files which include all the file names, in *fits[1] format, of flats in 1 band
- Command examples:
 - In a terminal:
 - Is NCvi0704*fits > targetk
 - In iraf:
 - files NCvi0704*fits > targetk
 - Modify with a text editor (gedit, emacs, vi ...)

 eskank@kankare-latitude:~/SN2012ej_20120907\$ Edit View Search Terminal Help eskank@kankare-latitude:~/SN2012ej_20120907\$ gethead OBJECT EXPTIME FILT2 * NCvi070016.fits skyflat 1 3 Ks NCvi070017.fits skyflat 2 3 Ks NCvi070018.fits skyflat 3 3 Ks NCvi070021.fits skyflat 4 3 Ks NCvi070022.fits skyflat 6 3 Ks NCvi070022.fits skyflat 1 3 Ks NCvi070053.fits skyflat 2 3 Ks NCvi070054.fits skyflat 3 3 Ks NCvi070055.fits skyflat 3 3 Ks NCvi070055.fits skyflat 3 3 Ks NCvi070055.fits skyflat 3 3 Ks NCvi070057.fits skyflat 4 3 Ks NCvi070057.fits skyflat 3 3 Ks NCvi070058.fits skyflat 3 3 Ks NCvi0704059.fits skyflat 3 3 Ks NCvi0704059.fits skyflat 3 3 Ks NCvi070454.fits SN2012ej 1 30 Ks NCvi070438.fits SN2012ej 4 30 Ks NCvi070441.fits SN2012ej 5 30 Ks NCvi070443.fits SN2012ej 5 30 Ks NCvi070443.fits SN2012ej 7 30 Ks NCvi070443.fits SN2012ej 7 30 Ks NCvi070444.fits SN2012ej 9 30 Ks 		
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Reduction example 1: notcam package & mkflat task

- Differential (bright and faint) flats with the same exposure time required
- Start iraf, open ds9, go to your data folder
- Download the bad pixel mask bad_zero_sci.fits www.not.iac.es/instruments/notcam/badmask/
- Uncompress and move to the data folder
- Open notcam package in iraf by typing notcam
- Select mkflat task by typing epar mkflat and set parameters, execute with :go

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NCvi070054.fits skyflat 2 3 Ks	
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NCvi070058.fits skyflat 6 3 Ks	
NCvi070059.fits skyflat 7 3 Ks	
NCvi070060.fits skyflat 8 3 Ks	
NCvi070437.fits SN2012ej 1 30 Ks	
NCvi070438.fits SN2012ej 2 30 Ks	
NCvi070439.fits SN2012ej 3 30 Ks	
NCvi070440.fits SN2012ej 4 30 Ks	
NCvi070441.fits SN2012ej 5 30 Ks	
NCvi070442.fits SN2012ej 6 30 Ks	
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eskank@kankare-latitude:~/SN2012ej_20120907\$ ls NCvi0700*fits > flatk	
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Space Telescope Tables Package TABLES Version 3.14	+ Image Reduction and Analysis Facility PACKAGE = notcam TASK = mkflat
Space Telescope Science Institute, Baltimore, Marylan Copyright (C) 2003 Association of Universities for Research in Astronomy, Inc.(AURA) See stsdas\$copyright.stsdas for terms of use. For help, send e-mail to help@stsci.edu	<pre>images = @flatk Input images [*] outim = flatk.fits Output file name badpix = bad_zero_sci.fits Input bad pixel image (zero pixels) camid = WF Camera (WF or HR) filtid = Ks Filter (Yn, J, H, K, Ks, or NOT number) suppress= yes Suppress dc-gradient structure? (datamax= INDEF) DATAMAX (if INDEF, saturation limit is used (autosea= ues) Auto-search the valid flatfields?</pre>
NOTCam scripts for quicklook reductions Version 2.5, Sep 2012	(flist =) (list =) (mode = ql)
phot mkflat reduce skysub bad mklincor reduce_bs	

Reduction example 1: master flat

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- Download the distortion correction model notcam.db for NOTCam WF camera from the NOT web pages: www.not.iac.es/instruments/notcam/distortion/
- Move notcam.db into the data folder
- Select reduce task by typing epar reduce and set parameters, execute with :go
- When prompted, select field star(s), common in the raw frames, close to the science target with key press 'a'. Finish with 'q'.
- If trim = no, FOV of the final reduced image will match the first frame.

● IRAF	
dif5574ba 21517. 1134.3 dif5574ca 21347. 1304.2	
Output image = flat_5574d, ncombine = 8 z1=0.742601 z2=1.261061 Normalized masterflat: flatk.fits - ready and displayed! Image mean stddev	IRAF IRAF Image Reduction and Analysis Facility PACKAGE = notcam TASK = reduce
<pre>flatk.fits[140:160,140:160] 1.050621 0.03505777 flatk.fits[350:370,350:370] 1.01404 0.03503872 flatk.fits[650:670,170:190] 0.990676 0.0337076 flatk.fits[700:720,370:390] 0.9903919 0.03873669 flatk.fits[00:620,600:620] 1.01906 0.04096627 flatk.fits[930:950,210:230] 0.9464503 0.04033541 flatk.fits[750:770,860:880] 1.001554 0.04181723 flatk.fits[350:370,810:830] 0.9856766 0.05816747 flatk.fits[350:370,810:830] 0.9971613 0.03434817</pre>	<pre>iimages = Bargetk Input images (or first image) nimages = 9 Number of input images output = SN2012ej_20120907_k.fits Output image flatfiel= flatk.fits Flat field image badpixma= bad_zero_sci.fits Bad pixel mask for fixpix (or INDEF) scale = add Scaling the sky (none, add, mult) combine = median Final imcombine (median or average) boundary= wrap Boundary type (nearest,constant,reflect,wrap) skip = yes Skip the first image when combining? trim = yes Trim the shifted images? badpixfi= no Interpolate over individual bad pixels? distcorr= yes Apply WF-cam distortion correction? destripe= no Remove stripes from bright stars? (imlist =) (mode = ql)</pre>
	EST-2 for HELP

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• IRAF	
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<pre>Oct 10 22:42: IMCOMBINE combine = median, scale = none, zero = median, weight = none reject = ccdclip, mclip = yes, nkeep = 1 rdnoise = 9.3, gain = 2.6425, snoise = 0., sigma = 3., hsigma = 3. blank = 0.</pre>	
Dutput image = SN2012ej_20120907_k.sky, ncombine = 9 z1=11424.39 zZ=17040.37 Sky subtraction done. Sky image displayed (NB! Includes the dark).	
2012 9 8 SAT 4:28:10.0 4:28:10.0 2012.68634 2456178.6862 2:27:12.7 Julian date is: 2456178.6862269 Pixel scale is: 0.234 Doing distortion correction Filter IDs: 0 207 Stop wheel: 15mm ring Using model: dist-k.dat please, wait	

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• IRAF	
NCvi070437.fits[1] NCvi070438.fits[1] NCvi070443.fits[1] NCvi070441.fits[1] NCvi070442.fits[1] NCvi070443.fits[1] NCvi070443.fits[1] NCvi070445.fits[1] Flat fielding done. Median background value: 14569.36 14446.47 14438.57 14430.02 14430.02 14430.66 14333.96	
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IRAF			
tmp\$im5574pq	145.871 (0.006)	474.580 (0.006)	2
tmp\$im5574sq tmp\$im5574sq	189.318 (0.010) 153.600 (0.006)	666.376 (0.011) 303.688 (0.006)	1 2
tmp\$im5574vq tmp\$im5574vq	360.254 (0.011) 324.455 (0.006)	675.027 (0.010) 311.962 (0.006)	1 2
tmp\$im5574yq tmp\$im5574yq	531.446 (0.010) 495.449 (0.007)	682.971 (0.011) 320.207 (0.006)	1 2
#Refcoords Reference tmp\$im25574zp tmp\$im25574zp	X-center Err 353.677 (0.012) 317.176 (0.008)	Y-center Err 845.518 (0.012) 482.851 (0.008)	Νum 1 2
#Shifts Image tmp\$im5574aq tmp\$im5574gq tmp\$im5574gq tmp\$im5574gq tmp\$im5574mq tmp\$im5574mq tmp\$im5574sq tmp\$im5574sq tmp\$im5574sq	X-shift Err -0.000 (0.010) -169.863 (0.011) -163.570 (0.011) 7.516 (0.009) 178.629 (0.009) 171.411 (0.010) 163.967 (0.009) -6.929 (0.010) -178.021 (0.010)	Y-shift Err 0.000 (0.010) -7.791 (0.011) -178.500 (0.011) -169.984 (0.010) -161.496 (0.010) 8.326 (0.009) 179.153 (0.010) 170.690 (0.010) 162.596 (0.010)	N Internal 2 (0.000,0.000) 2 (0.256,0.034) 1 (INDEF,INDEF) 2 (0.132,0.000) 2 (0.132,0.000) 2 (0.103,0.054) 2 (0.391,0.000) 2 (0.351,0.201) 2 (0.252,0.060)
#Trim_Section = [180	:845,181:845]		
# Shifting images:			
# Trimming images:	corrected section =	[180:845,181:845]	
Skipping the first	image when combining		
<pre>Oct 10 22:42: IMCOMB combine = median, : reject = minmax, n blank = 0. statsec = Oct 10 2: tmp\$im557; tmp\$im57; tmp\$im557; tmp\$im557; tmp\$im57; tmp\$im57;</pre>	INE scale = none, zero = low = 1, nhigh = 0 2:42 es Median Zero 4dq 12.418 0. 4gq 14.684 -2.2661 4jq 10.507 1.9107 4mq 18.688 -6.2697 4pq 5.8979 6.5202 4sq 5.9774 6.4007 4vq -5.6366 18.055 4gq 15.462 -3.0434	median, weight = nor	le
Output image = SN2 The images are flat aligned, shif	012ej_20120907_k, ncc field corrected, sky ted and combined.	ombine = 8 subtracted	
SN2012ej_20120907_k0 SN2012ej_20120907_k0 SN2012ej_20120907_k0 SN2012ej_20120907_k0 SN2012ej_20120907_k0 SN2012ej_20120907_k0 SN2012ej_20120907_k0 SN2012ej_20120907_k0 SN2012ej_20120907_k0 z1=36.33753_z2=2126.; Image SN2012ej_20120 notcam>	01 02 03 04 05 06 07 08 09 2785 0907_k is displayed.		
IRAF	SAOImage ds	9	



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• IRAF			e	0
tmp\$im5574sq tmp\$im5574sq	189.318 (0.010) 153.600 (0.006)	666.376 (0.011) 303.688 (0.006)	1 2	
tmp\$im5574vq tmp\$im5574vq	360.254 (0.011) 324.455 (0.006)	675.027 (0.010) 311.962 (0.006)	1 2	
tmp\$im5574yq tmp\$im5574yq	531.446 (0.010) 495.449 (0.007)	682.971 (0.011) 320.207 (0.006)	1 2	
#Refcoords Reference tmp\$im25574zp tmp\$im25574zp	X-center Err 353.677 (0.012) 317.176 (0.008)	Y-center Err 845.518 (0.012) 482.851 (0.008)	Num 1 2	
#Shifts Image tmp\$im5574aq tmp\$im5574dq tmp\$im5574gq tmp\$im5574gq tmp\$im5574pq tmp\$im5574pq tmp\$im5574pq tmp\$im5574vq tmp\$im5574vq tmp\$im5574vq	X-shift Err -0.000 (0.010) -169.863 (0.011) -153.570 (0.011) 7.516 (0.009) 178.629 (0.009) 171.411 (0.010) 163.967 (0.009) -6.929 (0.010) -178.021 (0.010)	Y-shift Err 0.000 (0.010) -7.791 (0.011) -178.500 (0.011) -163.984 (0.010) -161.496 (0.010) 8.326 (0.009) 179.153 (0.010) 170.690 (0.010) 162.596 (0.010)	N Internal 2 (0.000,0.00 2 (0.256,0.00 1 (INDEF,INDE 2 (0.132,0.00 2 (0.107,0.00 2 (0.103,0.00 2 (0.391,0.00 2 (0.351,0.20 2 (0.252,0.00	1 34) EF) 30) 85) 54) 30) 30)
#Trim_Section = [180:	845,181:845]			
# Shifting images:				
# Trimming images: c	orrected section =	[180:845,181:845]		
Skipping the first i	mage when combining			
Uct 10 22:42: IMCUMB: combine = medLamB: reject = minmax, nl blank = 0. statsec = Oct 10 22 Image tmp%im5574 tmp%im5574 tmp%im5574 tmp%im5574 tmp%im5574	NE ccale = none, zero = ow = 1, nhigh = 0 ::42 's Median Zero dq 12.418 0. gq 14.684 -2.2661 jq 10.507 1.9107 mq 18.688 -6.2697 pq 5.8979 6.5202 sq 5.9774 6.4407 vq -5.6366 18.055 yq 15.462 -3.0434	median, weight = nor	ne	
Output image = SN20 The images are flatf aligned, shift	12ej_20120907_k, ncc ield corrected, sky ed and combined.	ombine = 8 subtracted		
SN2012ej_20120907_k00 SN2012ej_20120907_k00 SN2012ej_20120907_k00 SN2012ej_20120907_k00 SN2012ej_20120907_k00 SN2012ej_20120907_k00 SN2012ej_20120907_k00 SN2012ej_20120907_k00 Z1=36.33753 z2=126.2 Image SN2012ej_20120 notcam> displ NCvi070 z1=9883.246 z2=19060. notcam> ■	1 2 3 4 5 6 7 8 9 7 785 907_k is displayed. 437.fits[1] 1 94			
	SAOImage ds	9		



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Reduction example 1: output files



[SAOImage ds9]

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- Example data available in /course_data/nir/SN2012ej_20121020.tar.gz
- Repeat the previous steps to create master flat(s) if not carried out yet
- Download and move the notcam.db file to the data folder as before: www.not.iac.es/instruments/notcam/distortion/
- Select reduce_bs task by typing epar reduce_bs and set parameters, execute with :go
- When prompted, select field star(s), common in the raw frames, close to the science target with key press 'a'. Finish with 'q'. Repeat with the sky (OFF) field.

• IRAF		
Image Redu PACKAGE = notcam TASK = reduce_bs	I R A F ction and Analysis Facility	
<pre>iimages = ■ @targetk nimages = 18 output = SN2012ej_20121020_k. flatfiel= flatk.fits badpixma= bad_zero_sci.fits scale = add combine = median boundary= wrap skip = yes trim = yes badpixfi= no distcorr= yes destripe= no (imlist = (mode = ql</pre>	 Input images (or first image) Total number of input images Flat field image Bad pixel mask image (or INDEF) Scaling the sky (none,add,mult) Final imcombine (median or average) Boundary type (nearest,constant,reflect,wrage) Skip the first image when combining? Trim the shifted images? Fix bad pixels? Apply WF-cam distortion correction? Remove stripes from bright stars?) 	антинининининининининининининининининини
	ESC-? for HELP	

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• IRAF	
NCvj200680.fits[1] NCvj200682.fits[1] NCvj200684.fits[1] NCvj200684.fits[1] NCvj200684.fits[1] NCvj200690.fits[1] NCvj200692.fits[1] NCvj200694.fits[1] NCvj200694.fits[1] NCvj200683.fits[1] NCvj200683.fits[1] NCvj200687.fits[1] NCvj200687.fits[1] NCvj200693.fits[1] NCvj200693.fits[1] NCvj200693.fits[1] NCvj200693.fits[1] NCvj200697.fits[1] NCvj200697.fits[1] NCvj200697.fits[1] NCvj200697.fits[1] NCvj200697.fits[1] NCvj200697.fits[1] NCvj200697.fits[1] 9 DN target and 9 DFF t Flat fielding done. Median background value of 11285.82 11366.74 11387.04 11274.27 11320.48 11619.32 Making the sky image	arget images found. of OFF field images:
Oct 11 20:03: IMCOMBINE combine = median, scale reject = ccdclip, mclip rdnoise = 9.3, gain = 2. blank = 0.	= median, zero = median, weight = none = yes, nkeep = 1 6425, snoise = 0.001, sigma = 3., hsigma = 3.
Images M tmp\$im6033bg 1 tmp\$im6033fg 1 tmp\$im6033fg 1 tmp\$im6033rg 1 tmp\$im6033rg 1 tmp\$im6033rg 1 tmp\$im6033rg 1 tmp\$im6033rg 1 tmp\$im6033rd 1 tmp\$im6033h 1	tedian Scale Zero 1404. 1.000 0. 1403. 1.000 0. 1393. 1.001 0. 1397. 1.001 0. 1395. 1.001 0. 1408. 1.000 0. 1405. 1.000 9.77E-4 1.322. 1.001 0. 1405. 1.000 0.
Output image = SN2012ej z1=8605.723 z2=14285.74 The 18 raw images are sk Sky image displayed (NB! # ASTTIMES: Observatory pa # timezone = 0 # longitude = 17:52. #*YR MON DAY ZT 2012 10 21 SUN 4:51:08.0 Doing distortion correcti Filter IDs: 1007 207 Stc	20121020_k.sky, ncombine = 9 xy subtracted. Includes the dark). arameters for Roque de los Muchachos, La Palma 8 UT EPOCH JD LMST 0 4:51:08.0 2012.80411 2456221.7022 5:39:46.4 .on
using model: dist-k.dat .	piease, wait



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NCv.j200696.fits[1]
 NCv,j200681.fits[1]
  NCv,j200683.fits[1]
 NCvj200685.fits[1]
 NCv,j200687.fits[1]
 NCv,j200689.fits[1]
 NCvj200691.fits[1]
 NCvj200693.fits[1]
 NCvj200695.fits[1
 NCv,j200697.fits[1]
  9 ON target and 9 OFF target images found.
  Flat fielding done.
  Median background value of OFF field images:
 11285.82
 11527.82
 11366.74
 11496.29
 11387.04
 11274.27
 11424.75
 11320.48
 11619.32
  Making the sky image ...
 Oct 11 20:03: IMCOMBINE
   combine = median, scale = median, zero = median, weight = none
   reject = ccdclip, mclip = yes, nkeep = 1
rdnoise = 9.3, gain = 2.6425, snoise = 0.001, sigma = 3., hsigma = 3.
   blank = 0.
             Images
tmp$im6033bg
                            Median Scale
                                             Zero
                           11404.
                                    1.000
                                               0.
             tmp$im6033fg
                            11403.
                                    1.000
                                               Ο.
             tmp$im6033jg
                            11393.
                                    1.001
                                               Ο.
             tmp$im6033ng
                            11397.
                                    1.001
                                                Ο.
             tmp$im6033rg
                            11395.
                                    1.001
                                               Ο.
             tmp$im6033vg
                            11408.
                                    1.000
                                               0.
             tmp$im6033zg
                            11405.
                                    1.000 9.77E-4
             tmp$im6033dh
                                   1.001
                           11392
                                               Ο.
             tmp$im6033hh 11405. 1.000
                                               0.
   Output image = SN2012ej_20121020_k.sky, ncombine = 9
  z1=8605.723 z2=14285.74
  The 18 raw images are sky subtracted.
Sky image displayed (NB! Includes the dark).
  # ASTTIMES: Observatory parameters for Roque de los Muchachos, La Palma
          timezone = 0
          longitude = 17:52.8
 ##YR MON DAY
                          ΖT
                                     UT
                                              EPOCH
                                                                JD
                                                                        LMST
  2012 10 21 SUN 4:51:08.0 4:51:08.0 2012.80411 2456221.7022 5:39:46.4
  Doing distortion correction ...
  Filter IDs: 1007 207 Stop wheel: 15mm ring
  Using model: dist-k.dat ... please, wait ...
  WF-camera distortion correction done.
 z1=-539.3416 z2=83.39968
  Select stars for ON target image alignment.
  Press 'a' on each star selected. Press 'q' when finished.
  Log file tmp$coo6033tf open
            LINE COORDINATES
     COL
                                                          ENCLOSED GAUSSIAN DIRECT
       R
             MAG
                   FLUX
                             SKY
                                     PEAK
                                             Е
                                                 PA
          845.17 366.28 845.17
  366.28
    6.02
          12.19 132816. -271.5
                                  19035. 0.04 -17
                                                              2.55
                                                                        2.08
                                                                              2.01
  331.20
          482.26 331.20 482.26
    6.46
          11.17 341017. -272.6
                                   44810. 0.06 -54
                                                              2.62
                                                                        2.27
                                                                               2.15
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• IRAF							
Flat fielding done. Median background valu 11285.82 11527.82 11366.74 11496.29 11387.04 11274.27 11424.75 11320.48 11619.32 Making the sky image .	ue of OFF	field :	images	s:			
Oct 11 20:03: IMCOMBINE combine = median, sc reject = ccdclip, mcl rdnoise = 9.3, gain = blank = 0. Images tmp\$im6033bg tmp\$im6033rg tmp\$im6033rg tmp\$im6033rg tmp\$im6033zg tmp\$im6033cg tmp\$im6033ch tmp\$im6033ch	ile = med ip = yes 2.6425, Median (11404. (11404. (11393. (11395. (11395. (11408.) 11405.) 11405.	ian, zen , nkeep snoise Scale 1.000 1.001 1.001 1.001 1.000 1.000 1.000	°0 = n = 1 = 0.0 Z€	nediar 001, ≤ 0. 0. 0. 0. 0. 0. 0. 0. 0. 0. 0. 0. 0.	n, weight = nor ;igma = 3., hs:	ne igma = 3.	
Output image = SN2012 z1=8605.723 z2=14285.7 The 18 raw images are Sky image displayed (M # ASTIIMES: Observatory # longitude = 17: ##YR MON DAY 2012 10 21 SUN 4:51: Doing distortion corres Filter IDs: 1007 207 Using model: dist-K.de WF-camera distortion c z1=-539.3416 z2=83.3996 Select stars for ON ta Press 'a' on each star	ej_20121 sky sub B! Inclu paramet 52.8 ZT 8.0 4:5 ction Stop when correction 8 reget ima selecte	020_k.sk tracted des the ers for 1:08.0 2 el: 15mr ease, wa n done. ge aligr d. Press	<y, no<br="">dark) Roque 2012.8 n ring ait s 'q'</y,>	combir de l 200CH 30411 30 when	ne = 9 .os Muchachos, JD 2456221.7022 finished.	La Palma LMS 5:39:46.4	T 4
 Log file tmp\$coo6033tf # COL LINE CODR # R MAG FLUX 366.28 845.17 366.28 6.02 12.19 132816. 331.20 482.26 331.20 6.46 11.17 341017. z1=-558.549 z2=31.51565 Select stars for OFF t Press 'a' on each star	open INATES SKY 845.17 -271.5 482.26 -272.6 arget im selecte	PEAK 19035. 44810. age alig d. Press	E 0.04 0.06 gnment	PA -17 -54	ENCLOSED 2.55 2.62 finished.	GAUSSIAN 2.08 2.27	DIRECT 2.01 2.15
Log file tmp\$coo6033uf # COL LINE COORT # R MAG FLUX 422.11 586.28 422.11 6.13 12.76 78853. 615.78 322.56 615.78 7.17 10.96 413674.	open DINATES SKY 586.28 -328.3 322.56 -325.4	PEAK 11104. 50684.	E 0.10 0.03	PA 83 49	ENCLOSED 2.59 2.64	GAUSSIAN 2.11 2.49	DIRECT 2.04 2.39
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# Shifting images:
  # Trimming images: corrected section = [74:951,73:951]
   Skipping the first image when combining ...
  Oct 11 20:04: IMCOMBINE
    combine = median, scale = none, zero = mode, weight = none
    reject = minmax, nlow = 1, nhigh = 0
    blank = 0.
    statsec = Oct 11 20:04
                     Images
                                  Mode
                                            Zero
               tmp$im6033dg -11.074
                                              0
                                162.73
               tmp$im6033hg
                                         -173.8
               tmp$im60331g
                                75.025 -86.099
               tmp$im6033pg
                                176.18 -187.25
               tmp$im6033tg 60.055 -71.129
               tmp$im6033xg 183.57 -194.64
               tmp$im6033bh -121.78 110.71
               tmp$im6033fh -264.41 253.34
   Output image = SN2012ej_20121020_k, ncombine = 8
  Oct 11 20:04: IMCOMBINE
    combine = median, scale = none, zero = mode, weight = none
    reject = minmax, nlow = 1, nhigh = 0
    blank = 0.
   statsec = Oct 11 20:04
               Images
tmp$im6033cg
                                  Mode
                                            Zero
                                     0
                                               \cap
                                1.2138 -1.2138
               tmp$im6033gg
               tmp$im6033kg -14.215
                                         14.215
               tmp$im6033og -7.6886
                                        7.6886
               tmp$im6033sg -3.2584 3.2584
               tmp$im6033wg 2.3974 -2.3974
               tmp$im6033ah 2.2315 -2.2315
               tmp$im6033eh -16.216 16.216
               tmp$im6033ih
                                    Ο.
                                               0.
    Output image = SN2012e,j_20121020_k_OFF, ncombine = 9
   The images are flatfield corrected, sky subtracted,
          aligned, shifted and combined.
 SN2012ej_20121020_k001
SN2012ej_20121020_k002
SN2012ej_20121020_k002
SN2012ej_20121020_k003
  SN2012ej_20121020_k004
  SN2012ej_20121020_k005
  SN2012ej_20121020_k006
  SN2012ej_20121020_k007
 SN2012ej_20121020_K007
SN2012ej_20121020_K008
SN2012ej_20121020_K_DFF001
SN2012ej_20121020_K_DFF003
SN2012ej_20121020_K_DFF003
SN2012ej_20121020_K_DFF003
  SN2012e,j_20121020_k_OFF005
  SN2012ej_20121020_k_OFF006
 SN2012ej_20121020_k_DFF007
SN2012ej_20121020_k_DFF007
SN2012ej_20121020_k_DFF008
SN2012ej_20121020_k_DFF009
  z1=-42.11141 z2=82.70922
  Image SN2012ej_20121020_k_OFF is displayed in frame 3.
z1=-59.70681 z2=114.9561
   Image SN2012ej_20121020_k is displayed in frame 4.
  notcam> []
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SN2012ej_20121	020											•
File Edit View Go E	Bookmarks Help											
	• 👼 Home SN201	2ej_20121020									← → Q s	ear
🖉 DELLUTILITY												
Computer	NCvj200062.fits	NCvj200063.fits	NCvj200064.fits	NCvj200065.fits	NCvj200066.fits	NCvj200067.fits	NCvj200068.fits	NCvj200069.fits	NCvj200090.fits	NCvj200091.fits	NCvj200092.fits	
Home Desktop Documents	NCvj200093.fits	NCvj200094.fits	NCvj200095.fits	NCvj200096.fits	NCvj200097.fits	bad_zero_sci.fits	# Mod # Wed begin notcam.db	NCvj2 NCvj2 NCvj2 NCvj2 NCvi2	NCvj2 NCvj2 NCvj2 NCvj2 NCvj2 Flatk	flatk.fits	SN2012ej_ 20121020_k.sky.fits	
Music Pictures Videos	NCvj200680.fits	NCvj200681.fits	NCvj200682.fits	NCvj200683.fits	NCvj200684.fits	NCvj200685.fits	NCvj200686.fits	NCvj200687.fits	NCvj200688.fits	NCvj200689.fits	NCvj200690.fits	
Trash												
Browse Net	NCvj200691.hts	NCvj200692.fits	NCvJ200693.hts	NCvJ200694.hts	NCvJ200695.fits	NCvj200696.fits	NCvj200697.hts	SN2012eJ_ 20121020_k001.fits	SN2012eJ_ 20121020_k002.fits	SN2012eJ_ 20121020_k003.fits	SN2012eJ_ 20121020_k004.fits	
	SN2012ej_ 20121020_k005.fits	SN2012ej_ 20121020_k006.fits	SN2012ej_ 20121020_k007.fits	SN2012ej_ 20121020_k008.fits	SN2012ej_ 20121020_k.fits	SN2012ej_ 20121020_k_OFF. fits	SN2012ej_ 20121020_k_ OFF001.fits	SN2012ej_ 20121020_k_ OFF002.fits	SN2012ej_ 20121020_k_ OFF003.fits	SN2012ej_ 20121020_k_ OFF004.fits	SN2012ej_ 20121020_k_ OFF005.fits	
	SN2012ej_ 20121020_k_ OFF006.fits	SN2012ej_ 20121020_k_ OFF007.fits	SN2012ej_ 20121020_k_ OFF008.fits	SN2012ej_ 20121020_k_ OFF009.fits								

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SN2012ej_20121	020											•
File Edit View Go E	Bookmarks Help											
	• 👼 Home SN201	2ej_20121020									← → Q s	ear
🖉 DELLUTILITY												
Computer	NCvj200062.fits	NCvj200063.fits	NCvj200064.fits	NCvj200065.fits	NCvj200066.fits	NCvj200067.fits	NCvj200068.fits	NCvj200069.fits	NCvj200090.fits	NCvj200091.fits	NCvj200092.fits	
Home Desktop Documents	NCvj200093.fits	NCvj200094.fits	NCvj200095.fits	NCvj200096.fits	NCvj200097.fits	bad_zero_sci.fits	# Mod # Wed begin notcam.db	NCvj2 NCvj2 NCvj2 NCvj2 NCvi2	NCvj2 NCvj2 NCvj2 NCvj2 NCvj2 Flatk	flatk.fits	SN2012ej_ 20121020_k.sky.fits	
Music Pictures Videos	NCvj200680.fits	NCvj200681.fits	NCvj200682.fits	NCvj200683.fits	NCvj200684.fits	NCvj200685.fits	NCvj200686.fits	NCvj200687.fits	NCvj200688.fits	NCvj200689.fits	NCvj200690.fits	
Trash												
Browse Net	NCvj200691.hts	NCvj200692.fits	NCvJ200693.hts	NCvJ200694.hts	NCvJ200695.fits	NCvj200696.fits	NCvj200697.hts	SN2012eJ_ 20121020_k001.fits	SN2012eJ_ 20121020_k002.fits	SN2012eJ_ 20121020_k003.fits	SN2012eJ_ 20121020_k004.fits	
	SN2012ej_ 20121020_k005.fits	SN2012ej_ 20121020_k006.fits	SN2012ej_ 20121020_k007.fits	SN2012ej_ 20121020_k008.fits	SN2012ej_ 20121020_k.fits	SN2012ej_ 20121020_k_OFF. fits	SN2012ej_ 20121020_k_ OFF001.fits	SN2012ej_ 20121020_k_ OFF002.fits	SN2012ej_ 20121020_k_ OFF003.fits	SN2012ej_ 20121020_k_ OFF004.fits	SN2012ej_ 20121020_k_ OFF005.fits	
	SN2012ej_ 20121020_k_ OFF006.fits	SN2012ej_ 20121020_k_ OFF007.fits	SN2012ej_ 20121020_k_ OFF008.fits	SN2012ej_ 20121020_k_ OFF009.fits								

SAOImage ds9

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Photometry

- Most likely photometric standard star observations are not needed. 2MASS JHK magnitudes of field stars can be used most of the time to derive zero points and calibrate images.
- GAIA is a useful software for quick photometry.
- Couple of basic features:
 - Overplot 2MASS stars:
 Data-Servers → Catalogs → 2MASS
 - Aperture photometry: Image-Analysis → Aperture photometry → Results in magnitudes
 - Adjust if necessary: Semimajor axis, Annulus inner scale, Annulus outer scale, Sky estimator: mean → clipped mean
- Do not use non-linear field stars for photometry
- PSF photometry beyond the scope of this lecture

GAIA



Top 10 things to remember

1) Linearity <25000 counts - Check & adjust exposure times if necessary

- 2) In near-IR the sky background is always high Dither & sky-subtract
- 3) Overhead factor ~1.4 More with beamswitch
- 4) FOV 4' x 4' Use beam-switch for extended objects
- 5) Basic scripts for observing setup-ima , 9point , beamswitch
- 6) Target and sky observations require the same total exposure time
- 7) frame exposure mode reduces noise Use if possible
- 8) Dead column in the centre of the array Include a skew in the observing grid
- 9) Guide probe area limited Select guide area carefully
- 10) notcam package mkflat , reduce , reduce_bs